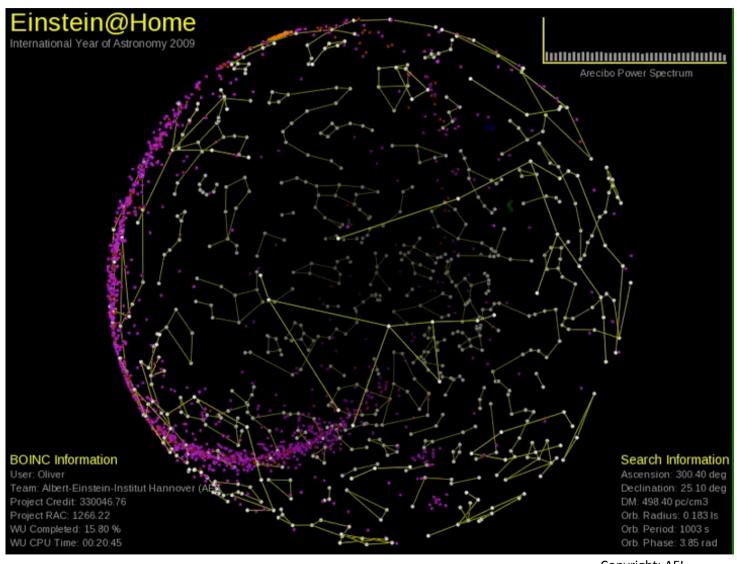


# Einstein@Home



Copyright: AEI

Bruce Allen
MPI for Gravitational Physics, Hannover



### **Einstein@Home Contributors**

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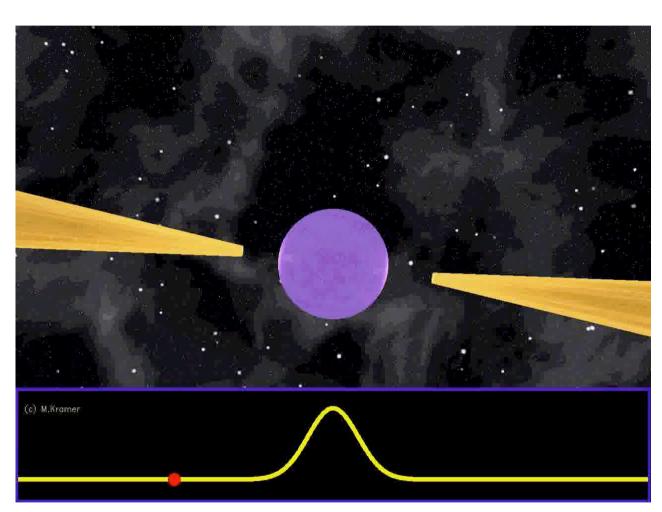
Karl Wette

Graham Woan

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### **Neutron Stars**





Copyright: Michael Kramer

- Discovered 1967
- About 2000 known
- Radius 8-12 km
- Up to 40 000 rpm
- 1 cm<sup>3</sup> weights 10<sup>15</sup> g (mountain)
- Very strong magnetic fields (10<sup>8</sup> - 10<sup>16</sup> Gauss)
- Emit (weak) radio waves, Xrays, gamma-rays. Should also emit (weak) gravitational waves



### The idea for Einstein@Home

- Lunch at the Caltech Athenaeum with Stuart Anderson (now head of computing for the LIGO laboratory)
- That morning we had both read this LA Times article

#### **Universe Wide Web**

August 19, 1999 | SARAH YANG | SPECIAL TO THE LOS ANGELES TIMES An ambitious project is enlisting the help of 1 million computer users worldwide to analyze radio signals from outer space in the search for extraterrestrial intelligence.

BERKELEY — Every day, every few seconds, some of the world's most powerful radio telescopes scan millions of channels in outer space in an effort to detect signs of extraterrestrial communication. No alien signals have been detected thus far. But is that because no one is out there contacting us, or are we just not listening hard enough?

Scientists are coming closer to the answer as legions of computer users around the world donate their PCs' otherwise idle time to help in the Search for Extraterrestrial Intelligence, or SETI. Scientists at UC Berkeley's Space Sciences Laboratory have been running the project, called SETI@home, for three months, in an ambitious hunt for signs of alien civilizations.

Earlier this week, the project passed a milestone by logging in its millionth participant.

"There are very few science projects out there that you can be a part of," said Dan Werthimer, chief scientist for SETI@home. "This is the first program where people can participate in a global science project. It's also a great way to get kids interested in science."







### Einstein@Home: a volunteer supercomputer

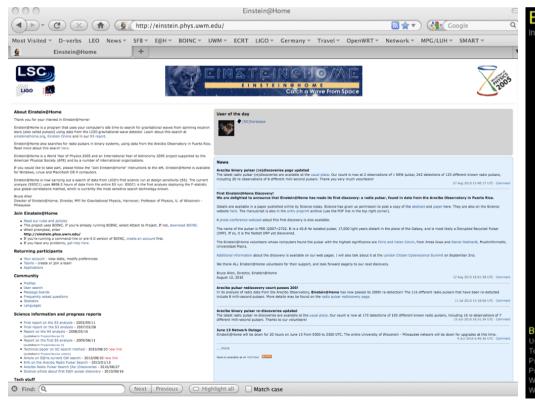
### **TARGET SOURCE:**

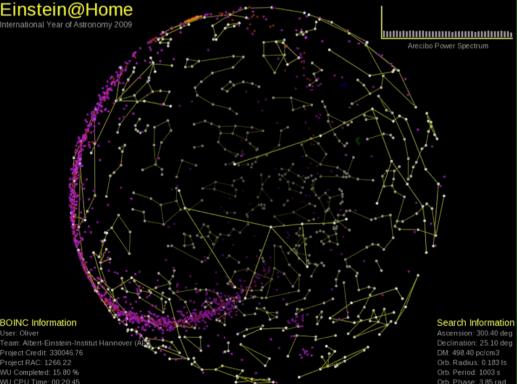
Rapidly-spinning neutron stars

LAUNCH: 2005 (Einstein Year)

### THREE DISTINCT SEARCHES:

- Gravitational Wave Data
- Radio Data (since 2009)
- Gamma Ray Data (since 2011)







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### **Einstein@Home Publications**

[Related papers on "methods" are not listed]

**Einstein@Home discovery of four young gammaray pulsars in Fermi LAT data**, H. J. Pletsch, L. Guillemot, B. Allen, et al., Astrophysical Journal Letters, 779, L11 (2013)

The Einstein@Home search for radio pulsars and PSR J2007+ 2722 discovery, B Allen, B Knispel, JM Cordes, JS Deneva, et al., Astrophysical Journal 773, 91 (2013)

**Einstein@Home all-sky search for periodic gravitational waves in LIGO S5 data**, LSC, Physical Review D 87, 042001, 2013.

Einstein@Home Discovery of 24 Pulsars in the Parkes Multi-beam Pulsar Survey, B Knispel, RP Eatough, H Kim, EF Keane, B Allen, et al., Astrophysical Journal 774, 93 (2013)

Arecibo PALFA Survey and Einstein@ Home: Binary
Pulsar Discovery by Volunteer Computing, B
Knispel Blazarus B Allen et al. The Astrophysical

Knispel, P Lazarus, B Allen, et al., The Astrophysical Journal Letters 732 (1), L1 (2011); P. Lazarus et al., Timing of a young mildly recycled pulsar with a massive white dwarf companion. MNRAS, 437, 1485-1494 (2014)

*Pulsar discovery by global volunteer computing*, B Knispel, B Allen, JM Cordes, JS Deneva, et al., Science 329 (5997), 1305-1305 (2010)

Einstein@ Home search for periodic gravitational waves in early S5 LIGO data, LSC, Physical Review D 80, 042003 (2009)

Einstein@ Home search for periodic gravitational waves in LIGO S4 data, LSC, Physical Review D 79, 022001 (2009)

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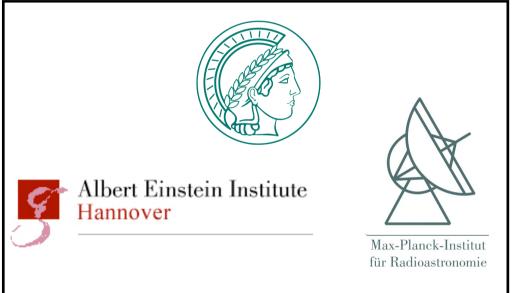
### **Einstein@Home Support**

# National Science Foundation through the UWM. Cornell.

through the UWM, Cornell, and UC Berkeley

# Max Planck Society through the AEI and MPIfR





- BOINC development (UCB)
- Gravitational wave searches (UWM)
- Server/data infrastructure (UWM)
- PALFA data preparation (Cornell)

- Programming
- Radio searches
- Gamma-ray searches
- Gravitational wave searches



### **Einstein@Home Computing**

- 3 different CPU search apps (radio, gamma-ray, gravitational wave)
- CUDA and OpenCL GPU versions for radio and gamma-ray search
- Android version of the radio search
- Past 9 years:377 000 volunteers193 UN countries
- Past two weeks:64 526 computers got work
- 10 393 active GPUs (NVIDIA/AMD/Intel)
- Currently 1232 Teraflops



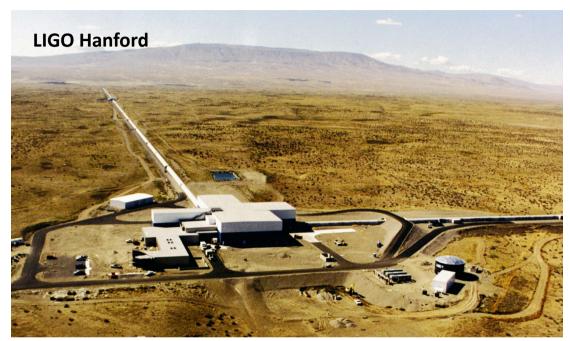


Status 28.9.2014, taken from http://einstein.phys.uwm.edu/server\_status.php

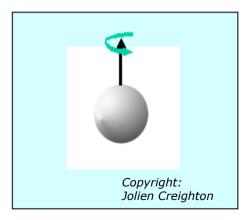
# Einstein@Home Gravitational Wave Searches



### **Gravitational Wave Detectors**







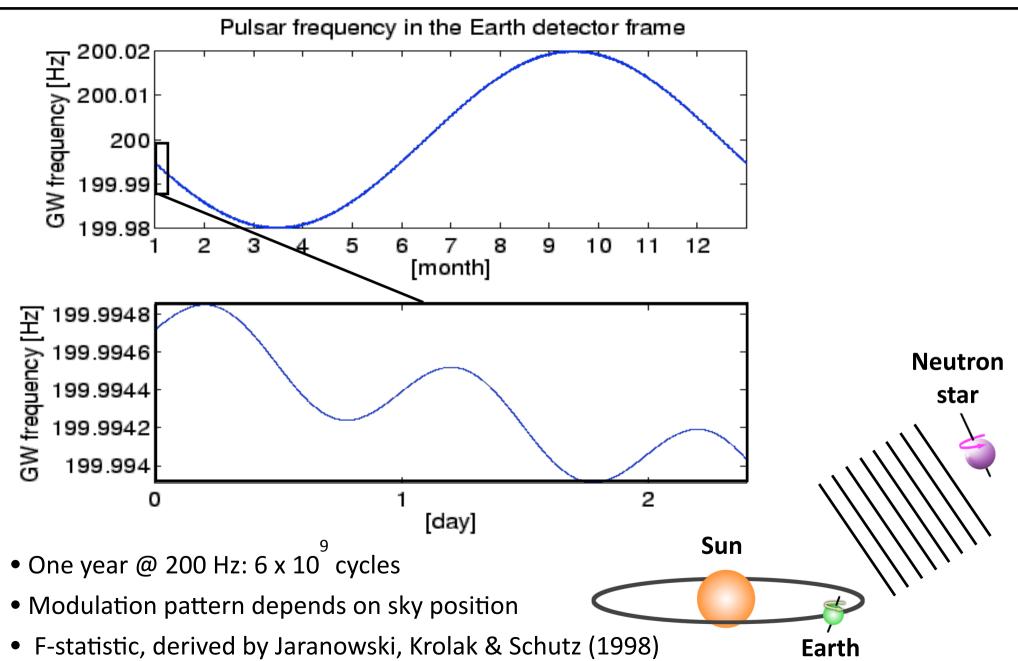
Mountain on a star

- Gravitational waves predicted by Einstein, 1916
- Observed indirectly, but not directly. Last remaining direct prediction of General Relativity
- Since mid-90s USA, France, Germany and UK building instruments to detect them
- In the USA, latest generation of the Laser Interferometer Gravitational-wave Observatories (LIGO) now starting commissioning: first extended science run expected in mid-2015

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## **GW Blind-Search Challenge**



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# Einstein@Home Radio Pulsar Search

## **Hunting for Radio Pulsars**



- Began in 2009
- Data from ongoing Arecibo PALFA survey, and archival Parkes PMPS survey
- New part of parameterspace: binaries with short periods
- Since mid-2010,
   Einstein@Home has
   discovered 48 new radio
   pulsars. These include a
   number of exotic binaries
   and milli-second pulsars.



Arecibo (Puerto Rico)

Copyright: Cornell University



Parkes (Australia)

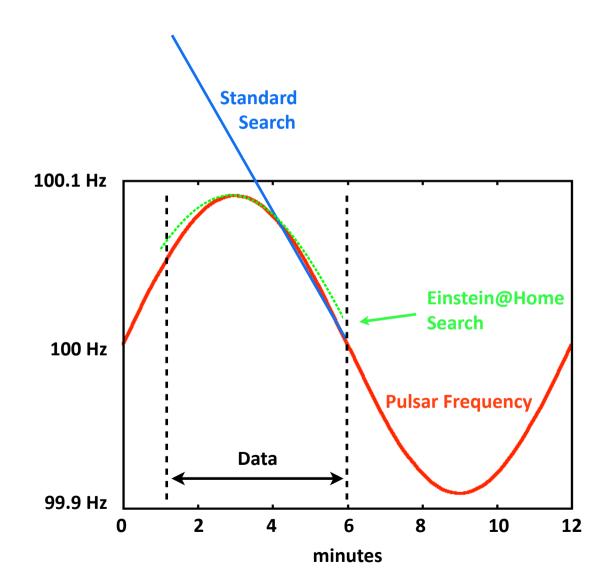
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### Radio Pulsar Search

- Einstein@Home does full orbital demodulation: sensitive to orbital periods as short as P<sub>orbit</sub> = 11 minutes
- "Standard search" looks for frequency changing linearly with time: loses sensitivity for P<sub>orbit</sub> < 50 minutes
- Special interest: binaries with two neutron stars (double neutron stars = DNS. Shortest known published P<sub>orbit</sub> are 2.5, 4.0, and 6.3 hours.
- Expectations for shortest-orbital period DNS in PALFA survey: 16 minutes (range 7 to 37 minutes)



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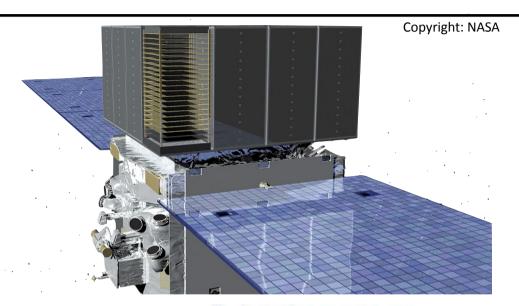
# Einstein@Home Search for Gamma Ray Pulsars

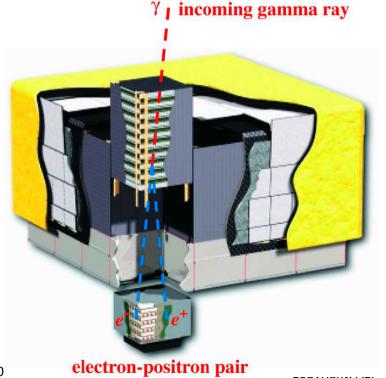


### Fermi Satellite Large Area Telescope (LAT)

- Launched June 2008
- Field of view: 20% of the sky, 20
   MeV 300 GeV
- Surveys the entire sky every 3 hours
- Public data set (150 million photons) with microsecond accuracy, ~1 degree sky-position
- Starting in 2011, gravitional wave search methods have found 15 new pulsars via gamma-ray emissions.
   Four of these from Einstein@Home
- Challenge: 8000 photons in 4 years:
   ~1 photon per 100000 revolutions

W. B. Atwood et al., *The large area telescope on the Fermi gamma-ray space telescope mission*, ApJ 697 (2009) 1071 doi:10.1088/0004-637X/697/2/1071



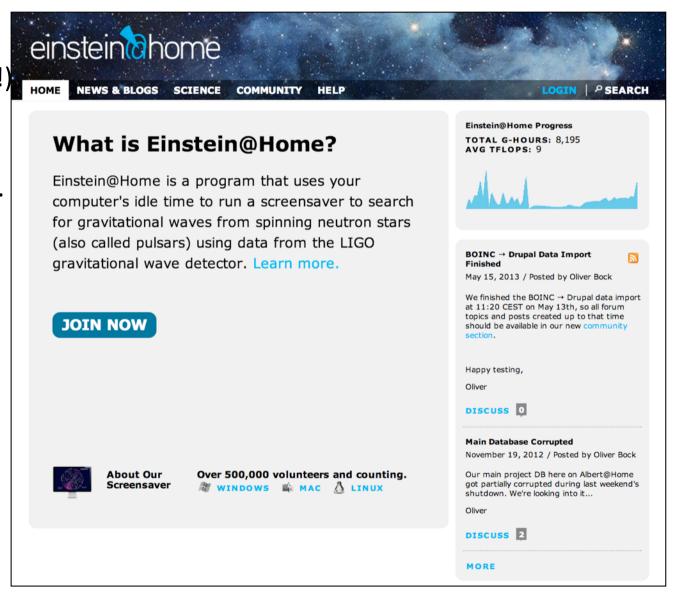


### **Outreach**



## **Web Pages**

- Hundreds of newspaper, magazine, radio and television stories (67 million Google hits!)
- Web site has community
  message boards, including a
  "Science" and "Tech" sections.
  Thousands of threads,
  including many detailed
  discussions with scientists.
- Web site also has "static" content about GWs, pulsars
- New Drupal-based web site in beta-test phase, hope to go public in some months. Will then add scientist and developer blogs, crowdsourcing



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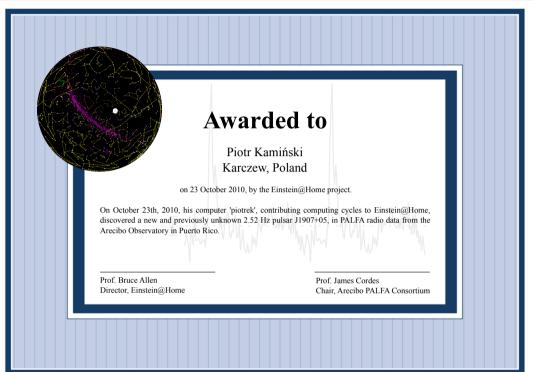
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## **Discovery Certificates**

- (Willing) volunteers are acknowledged by name in discovery papers
- They receive two formal framed "discovery certificates", one in English, and one in their native language
- Currently we have English (USA, Australia, South Africa, UK), German (Switzerland, Germany), Dutch, Russian, Ukranian, Japanese, Polish, Finnish, Slovenian, Czech, Hungarian, Turkish, French (Switzerland), Italian







### **Conclusions**





- Powerful tool for computationally-bound searches and a direct way for the public to contribute to scientific research.
- Gravitational wave searches will continue with improved methods and better data from Advanced LIGO
- Applying "GW technology" to radio and gamma-ray data is effective and will continue.
   Since mid-2010 we have found more than fifty new pulsars in radio data from Arecibo and Parkes, and gamma-ray data from Fermi LAT.
- Evolving with technology: support for NVIDIA and AMD/ATI GPUs and Android devices
- Future: crowdsourcing for post-processing, data storage AND computing (MeerKAT, SKA)



